

# Teleparallel Newton–Cartan gravity

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- ‘Non-relativistic’ limit: GR  $\rightarrow$  Newtonian gravity
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  - Galilei-relativistic spacetime geometry
  - Newtonian gravity described by curved connection

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  - Textbook formulation: linearised gravity
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- Geometric formulation:  $\text{GR} \xrightarrow{c \rightarrow \infty} \text{Newton–Cartan gravity}$ 
  - Galilei-relativistic spacetime geometry
  - Newtonian gravity described by curved connection
- Geometric description of the  $c \rightarrow \infty$  limit of teleparallel gravity?
  - Here: consider just TEGR
  - Read, Teh (2018): special case of our formalism; null-reduction instead of  $c \rightarrow \infty$

J Read, NJ Teh: *The teleparallel equivalent of Newton–Cartan gravity*, arXiv:1807.11779,  
Class. Quantum Gravity **35**, 18LT01 (2018)

# Outline

- 1 Recap of usual Newton–Cartan gravity
  - Galilei manifolds
  - Newton–Cartan gravity
  
- 2 Teleparallel Galilei connections
  - Bargmann structures
  - Teleparallel Newton–Cartan gravity
  
- 3 Relation to other theories
  - Teleparallel NC from TEGR
  - Recovering Newtonian gravity

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- This talk: mostly assume  $d\tau = 0$  (absolute time)

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- $\rightsquigarrow$  Galilei frame bundle  $G(M) \xrightarrow{\pi} M$ , principal Gal-bundle

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- *Newtonian connection*:  $T = 0$  and  $d\Omega = 0$



# Newton–Cartan gravity

## Axioms for Newton–Cartan gravity

- 1 Spacetime is a Galilei manifold  $(M, \tau, h)$  with absolute time, endowed with a Newtonian connection  $\tilde{\omega}$ ,
- 2 ideal clocks measure time as defined by  $\tau$ ,
- 3 free test particles move on timelike geodesics of  $\tilde{\omega}$ , i.e. timelike curves  $\gamma$  solving

$$\tilde{\nabla}_{\dot{\gamma}} \dot{\gamma} = 0, \quad (4)$$

- 4 the field equation

$$\tilde{R}_{\mu\nu} = 4\pi G \rho \tau_{\mu} \tau_{\nu} \quad (5)$$

holds, where  $\rho$  is the mass density.

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- Arises as formal  $c \rightarrow \infty$  limit of GR
- Coordinate formulation of Newtonian gravity can be recovered (up to possibility of non-absolute rotation)

# Teleparallel Galilei connections

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# ‘Teleparallelisation’ of NC gravity?

- ‘Teleparallelisation’ GR  $\rightarrow$  TEGR:
  - Lorentzian manifold  $(M, g)$ : unique torsion-free metric connection  $\overset{\text{L}}{\omega}$
  - $\implies$  General metric connection  $\overset{\text{L}}{\omega}$  determined by its torsion  $\overset{\text{L}}{T}$
  - $\rightsquigarrow$  Riemannian curvature  $\overset{\text{L}}{R}$  expressible purely in terms of  $\overset{\text{L}}{\omega}, \overset{\text{L}}{T}, \overset{\text{L}}{R} = 0$
  - $\rightsquigarrow$  Reformulation of GR in terms of  $\overset{\text{L}}{\omega}, \overset{\text{L}}{T}$

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  - $\rightsquigarrow$  Reformulation of GR in terms of  $\overset{\text{L}}{\omega}, \overset{\text{L}}{T}$
- Analogue for NC gravity?
  - Problem: Galilei connection  $\omega$  not uniquely determined by  $T$ !

# Extending Galilei to Bargmann structures

- Bargmann group  $\text{Barg} = \text{Gal} \ltimes (\mathbb{R}^4 \times \text{U}(1))$
- Extend Galilei frame bundle to Barg-bundle  $B(M) = G(M) \times_{\text{Gal}} \text{Barg}$

‘Globalised’ construction from:

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- Pulled back to  $M$ : *extended coframe* ( $\mathbf{e}^t = \tau, \mathbf{e}^a, \mathbf{ia}$ )
- Transformation under local boost of frame  $v \rightarrow v - k^a \mathbf{e}_a$ :

$$\mathbf{a} \rightarrow \mathbf{a} + k_a \mathbf{e}^a + \frac{1}{2} |k|^2 \tau \quad (6)$$

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# Extended torsion

- Exterior covariant derivative of  $(\tau, e^a, ia)$ : *extended torsion*

$$d^\omega(\tau, e^a, ia) = (T^t, T^a, \mathbf{if}) \quad (7)$$

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- For general  $\omega$ : *Newton–Cartan contortion*

$$\Gamma_{\mu\nu}^\rho - \tilde{\Gamma}_{\mu\nu}^\rho = \frac{1}{2}T_{\mu\nu}^\rho - T_{(\mu\nu)}{}^\rho + \tau_{(\mu} f_{\nu)}{}^\rho =: K_{\mu\nu}^\rho \quad (8)$$

# Teleparallel Newton–Cartan gravity

## Axioms for teleparallel Newton–Cartan gravity

- 1 Spacetime is a Galilei manifold  $(M, \tau, h)$  with absolute time, endowed with a Bargmann structure and a flat Galilei connection  $\omega$ ,
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$$(\nabla_{\dot{\gamma}} \dot{\gamma})^\rho = K^\rho_{\mu\nu} \dot{\gamma}^\mu \dot{\gamma}^\nu, \quad (9)$$

- 4 the field equation

$$-D_\sigma K^\sigma_{AB} + D_A K^\mu_{\mu B} - K^\mu_{\sigma B} T^\sigma_{\mu A} + K^\mu_{\mu\sigma} K^\sigma_{AB} - K^\mu_{A\sigma} K^\sigma_{\mu B} = 4\pi G \rho \tau_A \tau_B \quad (10)$$

holds, where  $\rho$  is the mass density.

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- LHS of (10) is  $\tilde{R}_{AB} \rightsquigarrow$  equivalent to usual NC gravity

# Relation to other theories

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$$E_{\mu}^0 = c\tau_{\mu} + c^{-1}a_{\mu} + \mathcal{O}(c^{-3}),$$

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$$\overset{\perp}{\omega}{}^0_0 = 0, \quad \overset{\perp}{\omega}{}^a_0 = c^{-1}\omega^a + \mathcal{O}(c^{-3}), \quad (12a)$$

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- Field equation of TEGR  $\rightsquigarrow$  trace-reverse  $\xrightarrow{c \rightarrow \infty}$  teleparallel NC field eq.!

# Recovering Newtonian gravity

- ‘Gauge fix’ to  $T^a_{bc} = 0$  [▶ Details](#)

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- For  $[e_a, v] = 0$ : equation of motion becomes

$$\ddot{\gamma}^a + \omega_c^a{}_b \dot{\gamma}^c \dot{\gamma}^b = -\partial^a \phi \quad (14)$$

# Conclusion

## Summary

- Bargmann structure formalism for NC geometry  $\rightsquigarrow$  natural notion of teleparallel Galilei connections
- Teleparallel formulation of NC gravity
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## Outlook

- TNC generalisation ( $d\tau \neq 0$ )?  $\rightsquigarrow$  probably needs different symmetry algebra ('TNC type II')
- Action formulation?
- Modified teleparallel gravity?
- Theories with more general geometry?
- 'Covariant' post-Newtonian limit of (modified) teleparallel gravity?

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**Many thanks for your attention!**

# Appendix: details

## 4 Details on recovery of Newtonian gravity

# Gauge fixing the purely spatial torsion

- Purely spatial part of field equation:  $\tilde{R}_{ab} = 0 \xrightarrow{3D}$  spatial metric flat
- $\implies$  We may assume

$$T^a_{bc} = 0 \tag{15}$$

consistently with flatness

◀ Back



# Trautman's 'absolute rotation' condition

- Usual NC: To recover Newtonian gravity, assume  $\tilde{R}^{ab}{}_{\mu\nu} = 0$
- $\iff \exists$  rigid, non-rotating frames
- Such frames in teleparallel NC:

$$\omega_{(ab)} = T_{(ab)t} \tag{16a}$$

$$\omega_{[ab]} = \frac{1}{2}f_{ab} \tag{16b}$$

◀ Back